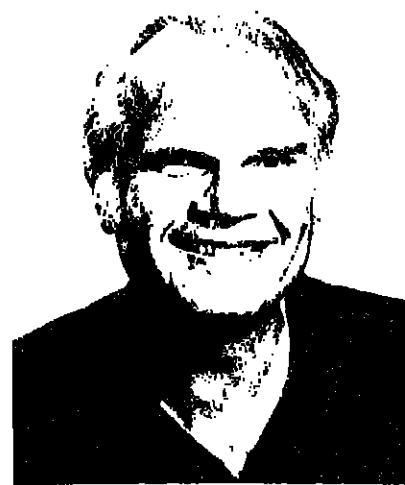


Eos

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John M. Wilcox 1925-1983



John Marsh Wilcox, Director of the Stanford Solar Observatory, died while swimming in surf in the Sea of Cortez near Puerto Penasco, Mexico, on October 14, 1983. He was 58 years old.

A solar physicist specializing in the study of magnetic fields at the solar surface and in the interplanetary medium, Wilcox authored or coauthored over 100 scientific publications. Over the past decade, working with various associates, Wilcox found that there is a yet-unexplained connection between the sector boundaries of the interplanetary magnetic

field and the axes of strong vorticity in the circulation of the earth's atmosphere at the lower boundary of the stratosphere. The effect, though small and of changeable magnitude, is of great theoretical interest.

Born in 1925, John Wilcox received his undergraduate education at Iowa State, and later obtained his Ph.D. in Physics from U.C., Berkeley in 1954. Professor Wilcox, in his early years, worked with Hannes Alfvén, helping with the first laboratory observations of Alfvén waves. He later turned his attention to the Sun and interplanetary medium and his influence upon the earth. He pioneered many studies of global solar structure, that is, viewing large scale solar phenomena (sectors, unipolar magnetic regions, etc.) and ignoring small scale structures (flares, spots, etc.), whose study was much in vogue in the 1960's. Working with Norman Ness, this approach bore fruit as the interplanetary medium, during most moderate levels of solar activity, was found to be ordered into large scale magnetic structures which John Wilcox called sectors. He worked with Robert Howard of Mount Wilson and showed that these structures were also found on the Sun.

In May 1981 Wilcox and his close associate Philip H. Scherrer discovered a method for identifying solar flares deemed likely to affect the earth's ionosphere, and through this to disturb global short-wave radio communications. These flares occurring in a southward pointing magnetic field tended to produce the ionospheric response, while those oppositely directed did not.

Wilcox will be remembered by many of his colleagues at the numerous meetings he attended, sporting a bright bowtie and a smile to match, discussing the latest findings on

sector related phenomena. These extend from cosmic rays, possibly to weather phenomena, which had been the object of many of John's recent studies.

The Stanford Solar Observatory was designed and built under Wilcox and Scherrer's direction and to specifications laid down by his research team. It enables extremely high accuracy observations of the sun's net magnetic field, observed in its total sunlight, as though it were a distant star. In collaboration with colleagues in the Crimean Astrophysical Observatory in the USSR, the Stanford group showed phase coherence in the velocity field oscillations.

Organizer and participant in national and international scientific meetings, Wilcox played an especially important role in the meetings of the American Geophysical Union. He was General Chairman of the Fall Meetings in 1978-79. He served as secretary of the Section on Interplanetary Physics in 1970-74, and in 1973-75 as an associate editor of the Journal of Geophysical Research.

He was active in meetings and symposia of the International Astronomical Union and the International Union of Geodesy and Geophysics. He also chaired the US participation in the US-USSR Bilateral Cooperation under the Environmental Working Group VIII from 1973 to 1979. He was an important and influential member of the Steering Committee for the International Symposium on Solar-Terrestrial Influences on Weather and Climate in Boulder in 1982. His contributions were carried out in the duties of a Member Representative to the University Corporation for Atmospheric Research from Stanford University.

He is remembered by his students at

Berkeley through the many pleasant evenings spent at "beer busts" at his house in the Berkeley hills. He tended to be a keen observer of life and related many of his observations during walks in the hills amidst the Eucalypti. He was proud of the fact that he had completed President Kennedy's 50 mile walk in the 1960's.

In addition to the great wealth of papers published, his influence will be most strongly felt by the many lives he touched. Wilcox was known to his scientific associates not only for his research accomplishments but for "his gentle manner of giving help to us when needed," as a group of Boulder scientists put it in a telegram of condolences at the occasion of his memorial service. He was a talented artist at the clarinet, and a person with a deep love for chamber music. At home, and wherever he traveled, he made time for musical events, especially when they involved small ensembles of great quality.

Wilcox is survived by his wife Ann, daughter Sharon, son David, and brothers William and Ralph.

Stanford University has created the John Marsh Wilcox Fund, to which donations may be made in his memory. The fund is dedicated to the acquisition of library materials for the Solar Observatory in the field of solar-terrestrial physics.

This tribute was written by Walter Orr Roberts, who is with the University Corporation for Atmospheric Research, Boulder, CO 80507, and Kenneth Howard Schatten, who is with the NASA Goddard Space Flight Center, Greenbelt, MD 20771.

News

IRAS: Taking Inventory

Even though its active life has ended, the Infrared Astronomical Satellite (IRAS) continues to make new discoveries as scientists at the Jet Propulsion Laboratory in Pasadena, Calif., pore over data from the orbiting satellite's year-long survey of the infrared universe. IRAS ran out of the cryogenic fluid that kept its telescopes and detectors cold enough for sensitive, infrared observations last November, thus ending the survey phase of the international project. Now begins the detailed study of IRAS data and the preparation of maps and catalogues of all infrared sources found by the satellite for distribution to the scientific community.

Among the discoveries already credited to IRAS are a ring of solid particles, possibly an evolving solar system, around the bright star Vega; six previously unknown comets; three narrow rings of dust lying within the plane of the solar system that may be the remains of asteroid or comet collisions; a small body designated 1983 TB that appears to be the burned-out cometary source of the annual Geminid meteor shower, and which passes closer to the sun at its perihelion than any planet or known asteroid; and many new and fascinating stellar and galactic sources.

Now IRAS scientists have added to the list another "proto-solar system" around the southern-hemisphere star Fomalhaut. A dimmer, cooler star than Vega, Fomalhaut shows an excess of infrared radiation surrounding it to a distance of approximately 80 astronomical units (sun-earth distances), just as Vega did. That excess is interpreted as the re-radiation of heat energy by solid particles, at least the size of grains of sand, surrounding the star, again like Vega. The particles could be as large as asteroids, but not as large as earth-type planets. Heat energy from the planets of our own solar system would be invisible to an infrared telescope observing from Fomalhaut because of their low ratio of surface area to volume.

Beyond the fact that there is something surrounding both stars, though, IRAS can't tell very much, and recent observations using the National Aeronautics and Space Administration's (NASA) Kuiper Airborne Observatory have provided conflicting evidence on whether the particles around Vega are sand-sized or whether they are smaller. Because Vega, Fomalhaut, and the sun are all in different stages of development, and because there are other stars in the IRAS survey with infrared excesses that may turn out to be best explained as Vega-type systems, some IRAS team members, like Program Scientist Nancy Boggess of NASA headquarters, hope that "after the data are processed well, we'll have a whole range of these systems in earlier or later stages of development to better understand planetary formation."

Asteroids are also turning up in the survey in big numbers—about 20,000 so far, which compares well with the 20,000 to 50,000 that IRAS team members expect to find once the full inventory is taken. The method of asteroid-hunting is to look for objects that appear in only one of the survey's four identical scans, then to see if it turns up in another location in another scan taken at a different time. There are several possible explanations

for an object that appears to move in this way—comet, asteroid, space debris, or even an artifact in the data—but an intriguing possibility is that one of them will turn out to be the 10th planet that some astronomers have theorized from observed perturbations in the motions of Uranus and Neptune. "There are many things that have appeared only once in the survey," says Boggess, "and the search is on."—TR

Permafrost Research

Among the highest priorities for permafrost research in the next 10 decades will be the improvement of detection and mapping techniques, the evaluation of engineering practices in frozen regions, and the establishment of a multidisciplinary program to observe the "active layer" just above the permafrost; these are the conclusions of a new report by the National Research Council's (NRC) Polar Research Board as part of its continuing effort to set guidelines for scientific research in NRC regions.

Like an earlier NRC report on Antarctic science (Eos, October 18, 1983, p. 356), the new study ranks the most pressing research priorities in an era of tight budgets. The 103-page study, entitled *Permafrost Research: An Assessment of Future Needs*, was prepared by the Polar Research Board's Committee on Permafrost, with the advice of scientists and engineers from state and federal agencies, private industry, and universities.

Permafrost is defined as ground that remains continuously below 0°C for a period of 2 years or more. Nearly a quarter of the earth's land surface qualifies under that definition—it occurs below the sea on continental shelves, in high mountains, and on northern continents (where it is as thick as 700 m in some places). It also became the object of great scientific interest and considerable frustration on the part of engineers during the construction of the Alaska oil pipeline in the 1970's.

The first priority set by the NRC committee would be to improve the detection and mapping of both land and subsurface permafrost and to gather data on its thermal composition, temperature, and rheological properties. Specifically, the report recommends that a "modest" drilling and temperature measurement program be established for different types of permafrost. Subsea drilling and probing programs could benefit from cooperation between government and industry in offshore oil-producing regions of Alaska, the report said.

The committee next calls for a refinement of existing models of heat and mass transport in permafrost, necessary for understanding its response to natural and human disturbances. A third priority is to establish a program to monitor existing structures like bridges and buildings constructed over permafrost to gain knowledge about engineering strengths or weaknesses. Identifying and classifying the ground ice contained in permafrost is the next item in the list, which is arranged "approximately in order of decreasing priority," although "all major recommendations merit continuing attention and support." This is followed by the

creation of a multidisciplinary program, involving both physical scientists and biologists, to observe the active layer above the permafrost that thaws during the summer.

The report further calls for an increased understanding of the physical and chemical properties of frozen ground—how, for example, it handles mechanical stresses and strains—and for improvements in permafrost technology, centralization of information banks, and international cooperation among researchers working in the field. Because study of permafrost is a good tool for monitoring climatic change, the committee also recommends an expansion of the network of meteorological stations in permafrost regions.

The National Research Council is the research arm of the National Academies of Sciences. Copies of the permafrost report are available in limited quantity from the Polar Research Board, 2101 Constitution Ave., N.W., Washington, DC 20418.—TR

South Pacific Mineral Cache

Recent deep-water sampling of mineral-rich crusts on the seafloor between the Hawaiian Islands and Samoa revealed deposits of cobalt, nickel, and manganese that are richer than previous samples, according to a team of scientists from the U.S. Geological Survey (USGS) and the Federal Republic of Germany aboard the research vessel S. P. Lee.

Thin plates of crust dredged from a seamount about 260 km northwest of Palmyra Atoll and Kingman Reef (U.S. territorial possessions roughly midway between Honolulu and American Samoa) had a cobalt concentration of 2.5%, or more than twice the concentration that earlier reconnaissance studies indicated would be found. The rock samples also contained 0.8% nickel and 32% manganese, compared to the estimated concentrations of 0.5% and 25%, respectively. The areas in which the deposits were found are part of the relatively unexplored ocean bottom included in the recently proclaimed 200-nautical-mile U.S. Exclusive Economic Zone (EEZ).

High concentrations of cobalt on seamounts northwest of Johnson Island were also confirmed by the marine geologists. Concentrations of cobalt, nickel, and manganese in the metallic crusts increase southward from Hawaii and reach a maximum on seamounts just north of the equator, they found. USGS scientists cautioned, however, that the relationships between metallic crust thickness, seamount type, geological structure, topography, and water depth are more complex than had been thought based on earlier work.

The U.S. has no domestic sources of cobalt, which is used in superalloys, magnetic materials, welding and hardfacing alloys, armor plate, cutting tools, and high-speed and wear-resistant steels. U.S. supplies currently are shipped mainly from central Africa.

Geophysicists

William P. Bishop was appointed deputy assistant administrator for satellites in the National Environmental Satellite, Data, and Information Service (NESDIS), part of the National Oceanic and Atmospheric

Administration. His previous positions include deputy director of the life sciences division of the National Aeronautics and Space Administration (NASA), deputy director of the NASA environmental observations division, and assistant director for waste management of the Nuclear Regulatory Commission's division of fuel cycle and materials safety.

Bruce A. Bell, professor of seismology and director of the seismographic stations at the University of California, Berkeley, was elected president of the California Academy of Sciences in the trustees. He also was elected vice chairman of the California Seismic Safety Commission.

Charles C. Counselman III, a professor in planetary science at the Massachusetts Institute of Technology, was awarded the 1983 Carl F. Corbitt Prize for his pioneering theoretical and practical work in the field of surveying and geodesy. The biennial, international award was founded in 1968 to promote scientific, technical, and commercial application activities in surveying. The award is named in honor of the man who was responsible for many significant advances in geodesic and photogrammetric instruments in the late 1890's and early 1900's. Counselman is the first American to win the prize.

In Memoriam

The following AGU members are recently deceased. Their primary section affiliation and year they joined AGU are given.

Robert C. Ganton, 63, Solar-Planetary Relations, 1959.

Herbert E. Hudson, 73, Hydrology, 1957.

John S. Fey, 85, Seismology, 1947.

Farwell G. Neff, 34, Hydrology (student member), 1983.

Rikio Maseki, 58, Geomagnetism and Paleomagnetism, 1967.

Congress Reconvenes

Bills on geophysics-related topics are among the items facing legislators as they reconvene this week for the second session of the 98th Congress. A status report on some of those bills appears in this issue in the new "Legislative Update" column, which will be published regularly in *Eos* to keep readers abreast of pending legislation relevant to geophysics.

Pertinent issues for the coming session include education, trade, and some environmental laws. Education, which promises to be a campaign issue in this election year, most recently became a topic of national attention

News (cont. on p. 26)

AGU MEMBERS

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Particles and Fields—Interplanetary Space

SOLAR WIND EFFECTS IN SOLAR AND INTERPLANETARY SPACE. The PRODUCTION OF GEOMAGNETIC INDUCTIONS IN THE SOLAR WIND BY PRIMARY GALACTIC PARTICLES—CHENG AND WANG. P. Michel and P. J. Beck. *Journal of Geophysical Research*, 88, 5000-5010, 1983. 11 pages, 1 figure. The authors present a model for the production of geomagnetic inductions in the solar wind by primary galactic particles. The model is based on the assumption that the solar wind is a plasma with a magnetic field. The authors calculate the induced magnetic field and compare it with the observed magnetic field. The results show that the model is in good agreement with the observations.

THE EFFECT OF THE SOLAR WIND ON THE PRODUCTION OF GEOMAGNETIC INDUCTIONS IN THE SOLAR WIND. P. Michel and P. J. Beck. *Journal of Geophysical Research*, 88, 5011-5020, 1983. 10 pages, 1 figure. The authors present a model for the production of geomagnetic inductions in the solar wind by primary galactic particles. The model is based on the assumption that the solar wind is a plasma with a magnetic field. The authors calculate the induced magnetic field and compare it with the observed magnetic field. The results show that the model is in good agreement with the observations.

THE EFFECT OF THE SOLAR WIND ON THE PRODUCTION OF GEOMAGNETIC INDUCTIONS IN THE SOLAR WIND. P. Michel and P. J. Beck. *Journal of Geophysical Research*, 88, 5021-5030, 1983. 10 pages, 1 figure. The authors present a model for the production of geomagnetic inductions in the solar wind by primary galactic particles. The model is based on the assumption that the solar wind is a plasma with a magnetic field. The authors calculate the induced magnetic field and compare it with the observed magnetic field. The results show that the model is in good agreement with the observations.

THE EFFECT OF THE SOLAR WIND ON THE PRODUCTION OF GEOMAGNETIC INDUCTIONS IN THE SOLAR WIND. P. Michel and P. J. Beck. *Journal of Geophysical Research*, 88, 5031-5040, 1983. 10 pages, 1 figure. The authors present a model for the production of geomagnetic inductions in the solar wind by primary galactic particles. The model is based on the assumption that the solar wind is a plasma with a magnetic field. The authors calculate the induced magnetic field and compare it with the observed magnetic field. The results show that the model is in good agreement with the observations.

THE EFFECT OF THE SOLAR WIND ON THE PRODUCTION OF GEOMAGNETIC INDUCTIONS IN THE SOLAR WIND. P. Michel and P. J. Beck. *Journal of Geophysical Research*, 88, 5041-5050, 1983. 10 pages, 1 figure. The authors present a model for the production of geomagnetic inductions in the solar wind by primary galactic particles. The model is based on the assumption that the solar wind is a plasma with a magnetic field. The authors calculate the induced magnetic field and compare it with the observed magnetic field. The results show that the model is in good agreement with the observations.

THE EFFECT OF THE SOLAR WIND ON THE PRODUCTION OF GEOMAGNETIC INDUCTIONS IN THE SOLAR WIND. P. Michel and P. J. Beck. *Journal of Geophysical Research*, 88, 5051-5060, 1983. 10 pages, 1 figure. The authors present a model for the production of geomagnetic inductions in the solar wind by primary galactic particles. The model is based on the assumption that the solar wind is a plasma with a magnetic field. The authors calculate the induced magnetic field and compare it with the observed magnetic field. The results show that the model is in good agreement with the observations.

THE EFFECT OF THE SOLAR WIND ON THE PRODUCTION OF GEOMAGNETIC INDUCTIONS IN THE SOLAR WIND. P. Michel and P. J. Beck. *Journal of Geophysical Research*, 88, 5061-5070, 1983. 10 pages, 1 figure. The authors present a model for the production of geomagnetic inductions in the solar wind by primary galactic particles. The model is based on the assumption that the solar wind is a plasma with a magnetic field. The authors calculate the induced magnetic field and compare it with the observed magnetic field. The results show that the model is in good agreement with the observations.

THE EFFECT OF THE SOLAR WIND ON THE PRODUCTION OF GEOMAGNETIC INDUCTIONS IN THE SOLAR WIND. P. Michel and P. J. Beck. *Journal of Geophysical Research*, 88, 5071-5080, 1983. 10 pages, 1 figure. The authors present a model for the production of geomagnetic inductions in the solar wind by primary galactic particles. The model is based on the assumption that the solar wind is a plasma with a magnetic field. The authors calculate the induced magnetic field and compare it with the observed magnetic field. The results show that the model is in good agreement with the observations.

THE EFFECT OF THE SOLAR WIND ON THE PRODUCTION OF GEOMAGNETIC INDUCTIONS IN THE SOLAR WIND. P. Michel and P. J. Beck. *Journal of Geophysical Research*, 88, 5081-5090, 1983. 10 pages, 1 figure. The authors present a model for the production of geomagnetic inductions in the solar wind by primary galactic particles. The model is based on the assumption that the solar wind is a plasma with a magnetic field. The authors calculate the induced magnetic field and compare it with the observed magnetic field. The results show that the model is in good agreement with the observations.

THE EFFECT OF THE SOLAR WIND ON THE PRODUCTION OF GEOMAGNETIC INDUCTIONS IN THE SOLAR WIND. P. Michel and P. J. Beck. *Journal of Geophysical Research*, 88, 5091-5100, 1983. 10 pages, 1 figure. The authors present a model for the production of geomagnetic inductions in the solar wind by primary galactic particles. The model is based on the assumption that the solar wind is a plasma with a magnetic field. The authors calculate the induced magnetic field and compare it with the observed magnetic field. The results show that the model is in good agreement with the observations.

THE EFFECT OF THE SOLAR WIND ON THE PRODUCTION OF GEOMAGNETIC INDUCTIONS IN THE SOLAR WIND. P. Michel and P. J. Beck. *Journal of Geophysical Research*, 88, 5101-5110, 1983. 10 pages, 1 figure. The authors present a model for the production of geomagnetic inductions in the solar wind by primary galactic particles. The model is based on the assumption that the solar wind is a plasma with a magnetic field. The authors calculate the induced magnetic field and compare it with the observed magnetic field. The results show that the model is in good agreement with the observations.

THE EFFECT OF THE SOLAR WIND ON THE PRODUCTION OF GEOMAGNETIC INDUCTIONS IN THE SOLAR WIND. P. Michel and P. J. Beck. *Journal of Geophysical Research*, 88, 5111-5120, 1983. 10 pages, 1 figure. The authors present a model for the production of geomagnetic inductions in the solar wind by primary galactic particles. The model is based on the assumption that the solar wind is a plasma with a magnetic field. The authors calculate the induced magnetic field and compare it with the observed magnetic field. The results show that the model is in good agreement with the observations.

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January 17, 1984

Particles and Fields—Ionosphere

THE EFFECT OF THE SOLAR WIND ON THE PRODUCTION OF GEOMAGNETIC INDUCTIONS IN THE SOLAR WIND. P. Michel and P. J. Beck. *Journal of Geophysical Research*, 88, 5121-5130, 1983. 10 pages, 1 figure. The authors present a model for the production of geomagnetic inductions in the solar wind by primary galactic particles. The model is based on the assumption that the solar wind is a plasma with a magnetic field. The authors calculate the induced magnetic field and compare it with the observed magnetic field. The results show that the model is in good agreement with the observations.

THE EFFECT OF THE SOLAR WIND ON THE PRODUCTION OF GEOMAGNETIC INDUCTIONS IN THE SOLAR WIND. P. Michel and P. J. Beck. *Journal of Geophysical Research*, 88, 5131-5140, 1983. 10 pages, 1 figure. The authors present a model for the production of geomagnetic inductions in the solar wind by primary galactic particles. The model is based on the assumption that the solar wind is a plasma with a magnetic field. The authors calculate the induced magnetic field and compare it with the observed magnetic field. The results show that the model is in good agreement with the observations.

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Meetings (cont. from p. 29)

NASA/GSFC, Greenbelt, MD 20771 will supply information on an interim basis.

Hydrology

Geochemistry/Water Quality of Catchments

A half-day session dealing with catchment hydrogeology will be convened under the joint sponsorship of the Volcanology, Geochemistry, and Petrology and Hydrology sections. The meeting will focus on the quantification of the hydrogeological response of catchments as inferred from field research, numerical programs, and theoretical calculations. Particular attention will be given to the problem of extrapolation of results from laboratory and small-scale field experiments on weathering reactions, soil formation, and cation exchange processes to the catchment scale in the face of the large spatial variability associated with both hydrological and chemical processes.

Abstracts should be submitted in standard AGU format to one of the session chairmen, Owen Birkner, USGS National Center, 12201 Sunrise Valley Dr., Reston, VA 20190 or George Hombberger, Department of Environmental Sciences, Clark Hall, University of Virginia, Charlottesville, VA 22903, by February 15. Please note that the original and two copies of the abstract should be sent to AGU by the February 23 abstract deadline.

Hilltop Hydrology

The Surface Runoff Committee of AGU's Hydrology Section is preparing a one-day session on Hilltop Hydrology, to be held at the 1983 Spring Meeting in Cincinnati. Eight speakers have already accepted an invitation to contribute their latest ideas and results.

Topics that will be presented include variable source area concepts, relations of the saturated zone upon infiltration, kinematic wave approximations to the hilltop flow processes, water budget models, and comparisons of various runoff model models.

If you plan to present a paper at this session, please send a copy of the abstract you are going to submit to AGU headquarters, no

later than February 1, to Peter Germann, Department of Environmental Science, Clark Hall, University of Virginia, Charlottesville, VA 22903, 801-924-0558.

Ocean Sciences

Ocean Drilling: Past and Future

Over the past 15 years, the *Glomar Challenger* has drilled at 624 sites throughout the world ocean and has contributed to major advances in our understanding of global tectonics, the evolution of continental margins and paleoenvironments, and the nature of oceanic crust. Fundamental geoscience questions have been both answered and generated, and hundreds of scientists from dozens of countries have contributed to the results of this remarkable project. At this juncture, between the conclusion of the Deep Sea Drilling Project and the initiation of the Ocean Drilling Program, it is appropriate that we assess the implications of scientific ocean drilling and focus on the opportunities afforded by future drilling from a vessel of enhanced capabilities.

Response of the Upper Ocean to Very Strong Wind II

The session is intended to be a follow-up to a similar special session held at the 1983 AGU Fall Meeting. Topics of interest for this session include tropical or mid-latitude storm forcing of upwelling, mixing and internal wave processes in the upper ocean as well as studies of thermodynamic or kinetic energy budgets related to such processes. Both theoretical and observational studies are encouraged. Convener for the session will be Peter G. Black, NOAA, Atlantic Oceanographic and Meteorological Laboratory, Miami, Florida.

SPR: Cosmic Rays

New Techniques and Applications of Geomagnetic Effects in Cosmic Ray Studies (Cosponsored by GP)

Recent improvements in cosmic ray detection from balloons and satellites, in geomagnetic field measurements from satellites, and

in accurate high-speed computation have enormously improved the accuracy and availability of cosmic ray trajectory calculations. These developments and their application to studies of primary cosmic ray isotopes, cosmic-ray-disturbed time variations, nucleides produced by cosmic rays in the atmosphere, and to studies of geomagnetism will be reviewed. Contributed papers are encouraged.

New Observations and Theories in Solar Flare Particle Acceleration (Cosponsored by SS)

This session is for display of material related to special sessions on flare particle acceleration and on present and future solar missions.

★ Plan to Attend ★

The AGU Chapman Conference on Collisionless Shock Waves in the Heliosphere

February 20-24
Silverado Country Club
Napa Valley, California

Convenor: R. G. Stone

This program will consist of 4 tutorials, 20 invited reviews, and 80 contributed papers. There will be oral sessions Monday through Friday and poster sessions on Tuesday and Wednesday afternoons. Subjects to be covered include why and where shocks form in the heliosphere; shock dynamics and evolution; shocks associated with solar activity; planetary bow shocks, corotation shocks, and shock-interactions; subcritical, supercritical, quasi-parallel, and quasi-perpendicular shocks; the foreshock; and dissipation mechanisms and particle acceleration mechanisms.

The registration fee is \$95, which includes an open reception and three group lunches. A banquet is scheduled for the evening of February 23, — cost \$15. Registration deadline is February 9, housing deadline January 31.

Call AGU for more information or to register.
800-424-2488 or 202-462-6903

This session features invited and contributed papers on recently discovered aspects of the production of energetic ions and electrons in solar flares.

Displays of Energetic Solar Phenomena (Poster Session) (Cosponsored by SS)

This session is for display of material related to special sessions on flare particle acceleration and on present and future solar missions.

Particles and Fields—
Ionosphere

5300 Ionospheric Disturbances
COMPOSITE EQUATORIAL SPREAD-VELOCITY SPECTRA FROM
HISD TO SHOT WAVELENGTHS
N. High and R. P. Thorne (Naval Research Laboratory, Code 4100, Washington, D. C. 20375)

In an effort to provide a full-spectrum classification of equatorial spread-F irregularities from large-scale ionospheric observations, we have analyzed a combination of ionospheric data which includes both the spread-F and the spread-F_{min} data. This analysis is presented in a paper which is available in the *Journal of Geophysical Research*, Vol. 89, No. 1, 1984, pp. 1-10.

5301 Ionospheric Disturbances
COMPOSITE EQUATORIAL SPREAD-VELOCITY SPECTRA FROM
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5302 Ionospheric Disturbances
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5305 Ionospheric Disturbances
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5306 Ionospheric Disturbances
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N. High and R. P. Thorne (Naval Research Laboratory, Code 4100, Washington, D. C. 20375)

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5307 Ionospheric Disturbances
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HISD TO SHOT WAVELENGTHS
N. High and R. P. Thorne (Naval Research Laboratory, Code 4100, Washington, D. C. 20375)

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hundred of. Large amplitude electrostatic waves that have a maximum near 100 Hz and an energetic ion population that shows evidence for both parallel and perpendicular acceleration to about 30 eV also accompany such events. The horizontal thermal ion drift (electric field) that accompanies these plasma phenomena has extremely large gradients over small spatial scales. Such a signature is consistent with a requirement for a large-scale magnetic field to be linked to the available field-aligned potential difference that may be applied to the plasma.

J. Geophys. Res., A, Paper 3A1582

5314 Total Electron Content
THE STUDY OF THE EFFECT OF SOLAR ECLIPSES ON THE IONOSPHERIC BASED ON SATELLITE RADIO OBSERVATIONS
Elizabeth A. Cohen (Physics Department, La Trobe University, Bundoora, Victoria 3083, Australia)

Baseline ionospheric observations during eclipses have provided such information on the behavior of the ionosphere. A combination of Faraday rotation and differential phase measurements as well as ionospheric data can provide information on the topside and bottomside ionospheric behavior during an eclipse as well as on the production rate. Recent attention has been directed towards a study of the dynamics of the F region during an eclipse and the relation between the amount of depletion in the total electron content and the per cent obscuration of the sun.

Further observations are also required in order to establish whether the ionosphere is responding to the total solar eclipse as predicted theoretically. In particular, it appears that a more sensitive technique such as differential Faraday rotation is required. The total solar eclipse of 11 June 1983 with its long totality time of 15 minutes over Indonesia and the Pacific will provide an ideal opportunity for further studies of these unanswered questions.

Rad. Sci., Paper 35195

5315 Wave propagation
DIRECT ACCESS TO PLASMA RESONANCE IN TWO-DIMENSIONAL RADIO EXPERIMENT
N. H. Rytov (113, University of Tromsø, 9001 Tromsø), T. T. T. T.

The concept of the linear conversion of radio waves into electrostatic waves is adapted to ionospheric radio heating experiments. It is identified an access to the plasma resonance through the radio window. By means of existing heating facilities large concentrations of electrostatic waves can be generated in the ionosphere. The e.s. waves are confined to a restricted region in space, horizontally displaced relative to the vertical direction. By means of a heating transmitter along the magnetic meridian, the displacement being southwards on the northern hemisphere. Ray tracing studies of the location of the resonance region relative to the transmitter are presented, and the horizontal displacement is estimated.

J. Geophys. Res., Space, Paper 3A1574

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Particles and Fields
Magnetosphere

5325 (How Shock Waves)
THE STRUCTURE OF THE SHOCK SURFACE IN THE HELIOSPHERE
H. D. Prater, E. H. Prater, J. A. Lockwood, and R. E. McPherson
Between 1 and 10 AU (Paper 3A1566)

A survey of Dayside Flux Transfer Events Observed by ISEE 1 and 2 Magnetometers (Paper 3A1524) Paul R. Goetz

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